



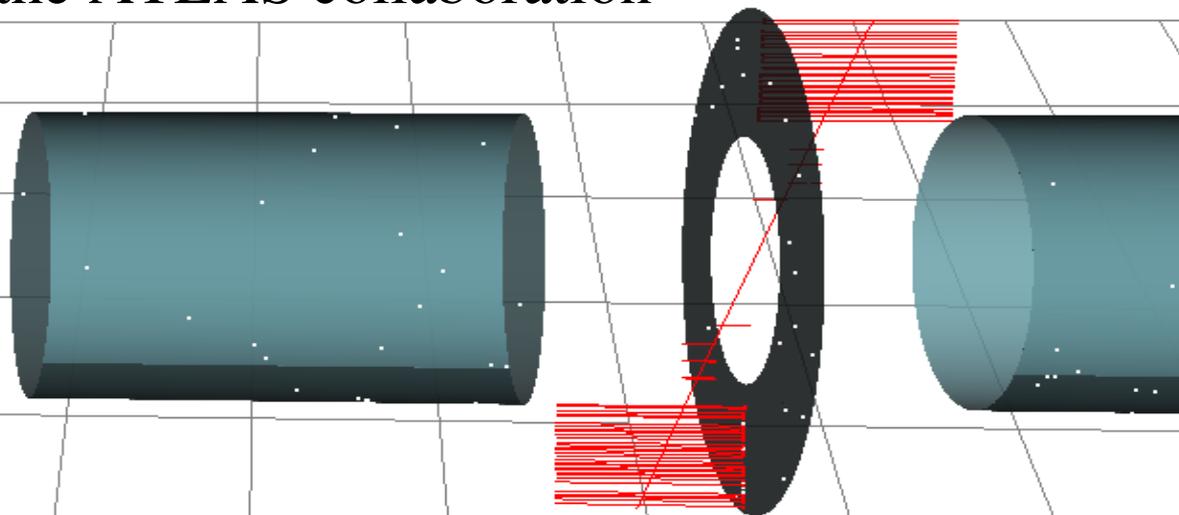
Alignment of the ATLAS Inner Detector Tracking System with Cosmic Ray Data

John Alison

University of Pennsylvania
on behalf of the ATLAS collaboration

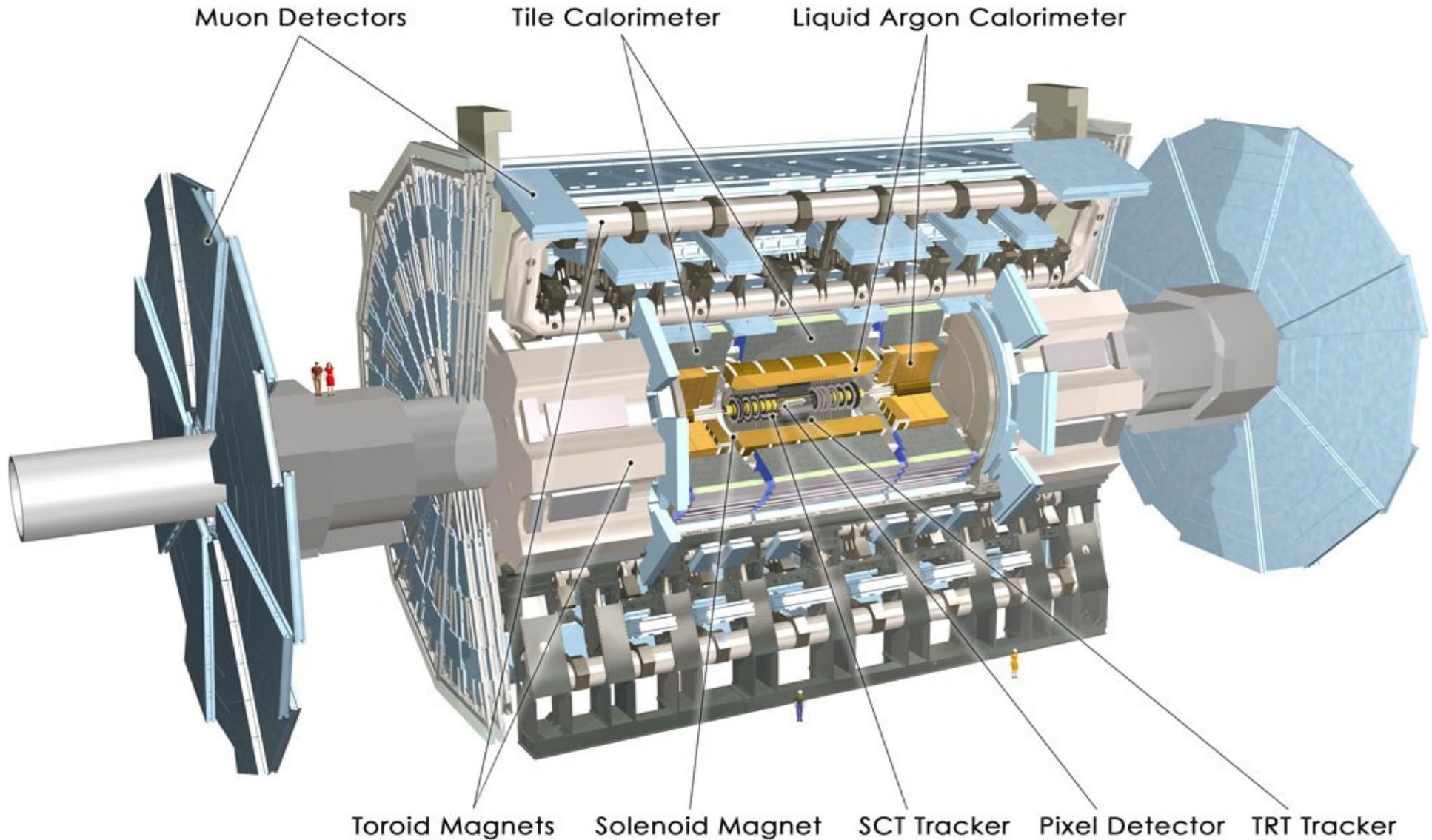
Outline:

- Introduction
- Track Based Alignment
- Results



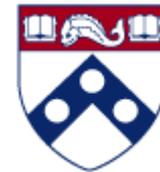


The ATLAS Detector

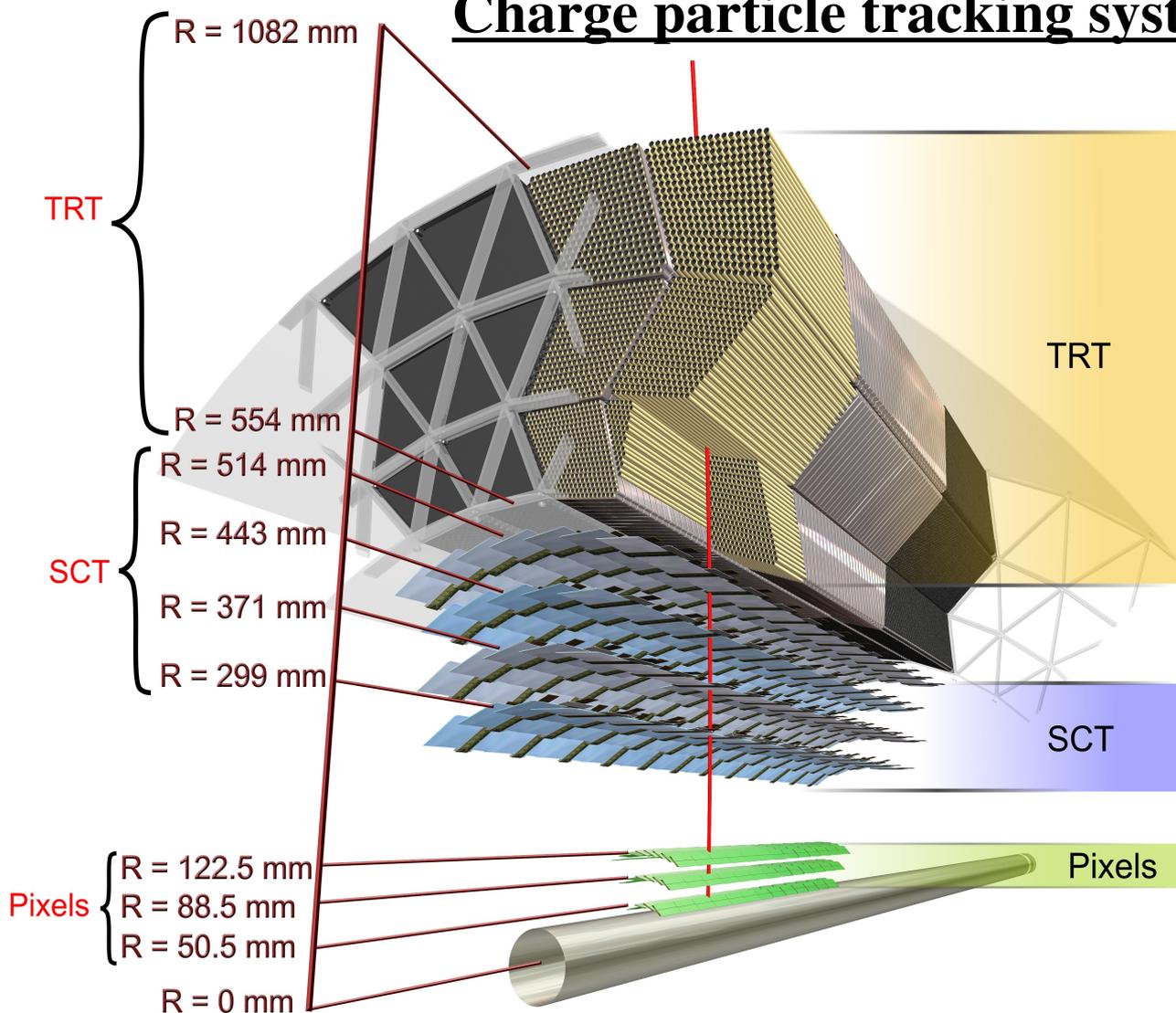




The ATLAS Inner Detector



Charge particle tracking system built on two technologies



Drift tubes:

~300,000 straw tubes
resolution $130 \mu\text{m}$ ($R\phi$)
Xe-CO₂-O₂
30 hits per track

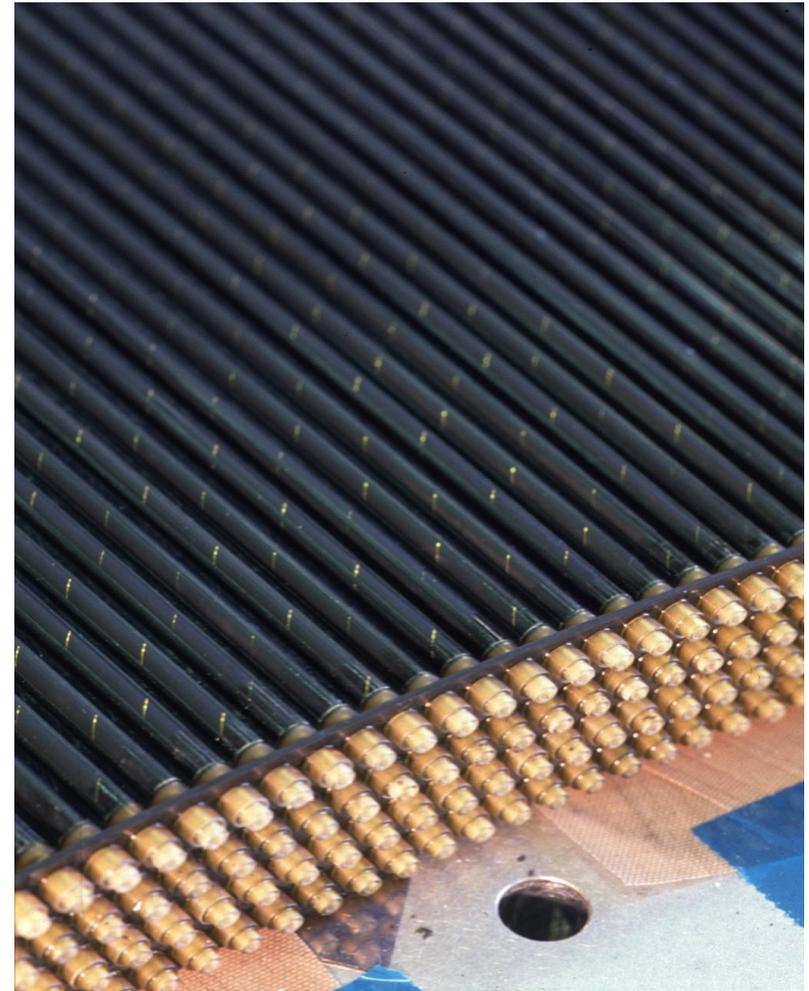
Silicon:

~ 6M Si strips
resolution: $17 \mu\text{m}$ ($R\phi$)
 $580 \mu\text{m}$ (Z)
4 barrel layers / 2 x 9 endcap disks

~ 80M Si pixels
resolution: $10 \mu\text{m}$ ($R\phi$)
 $115 \mu\text{m}$ (Z)
3 barrel layers/ 2 x 3 endcap disks

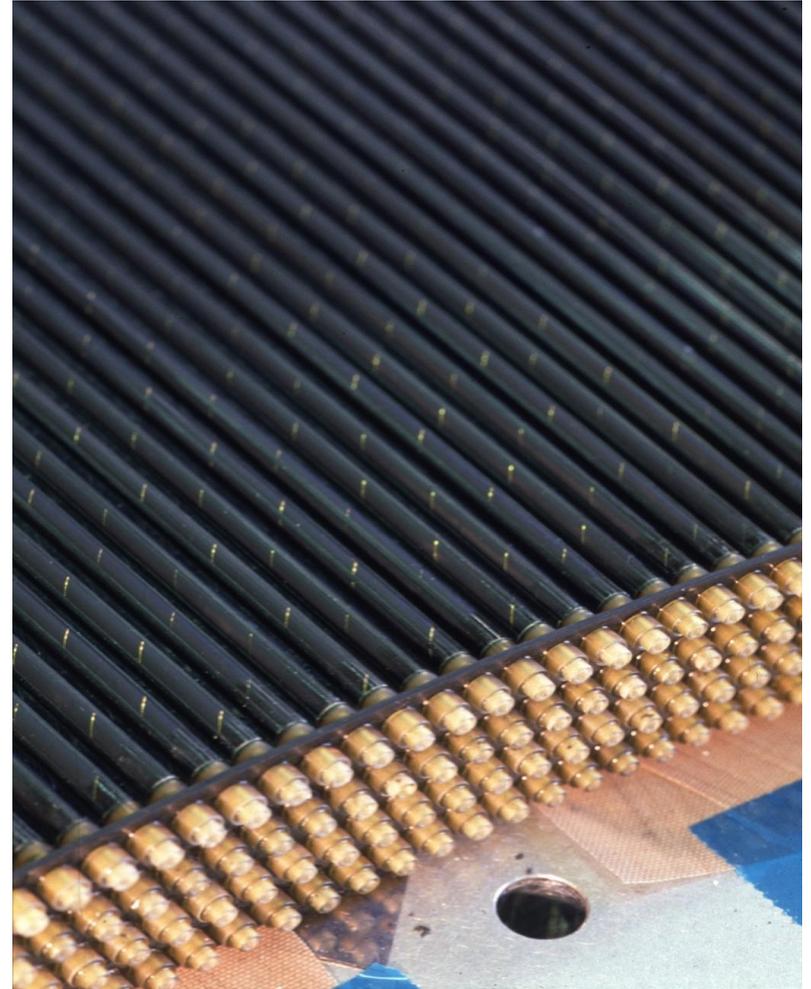


Construction highly modular



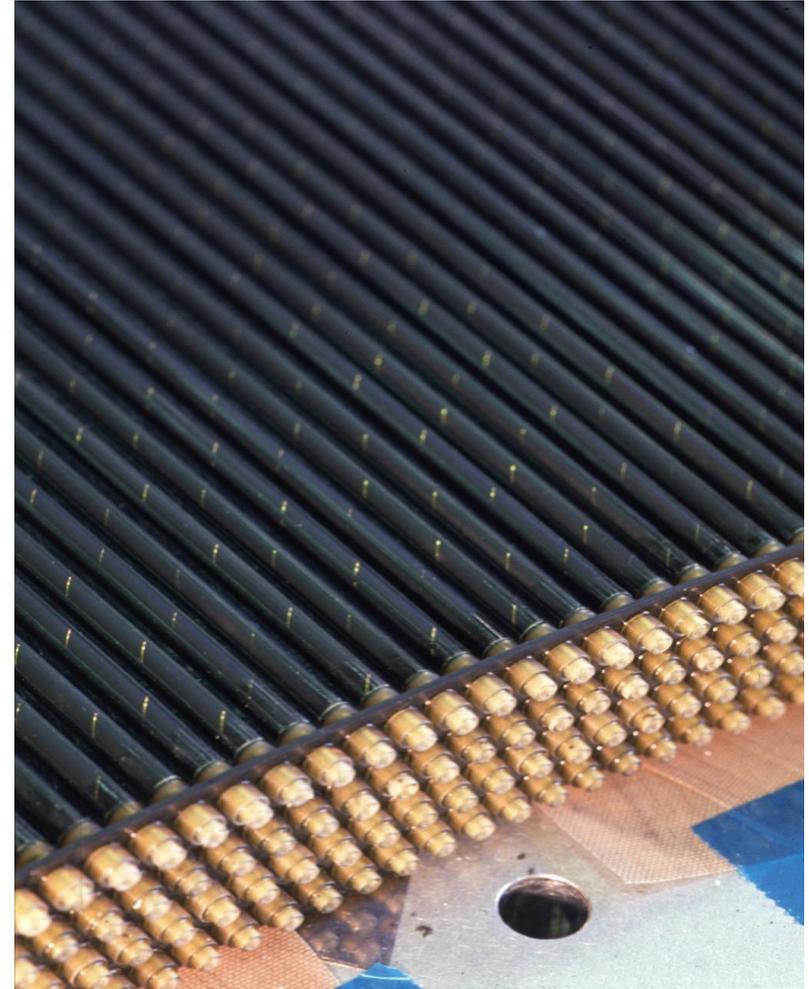
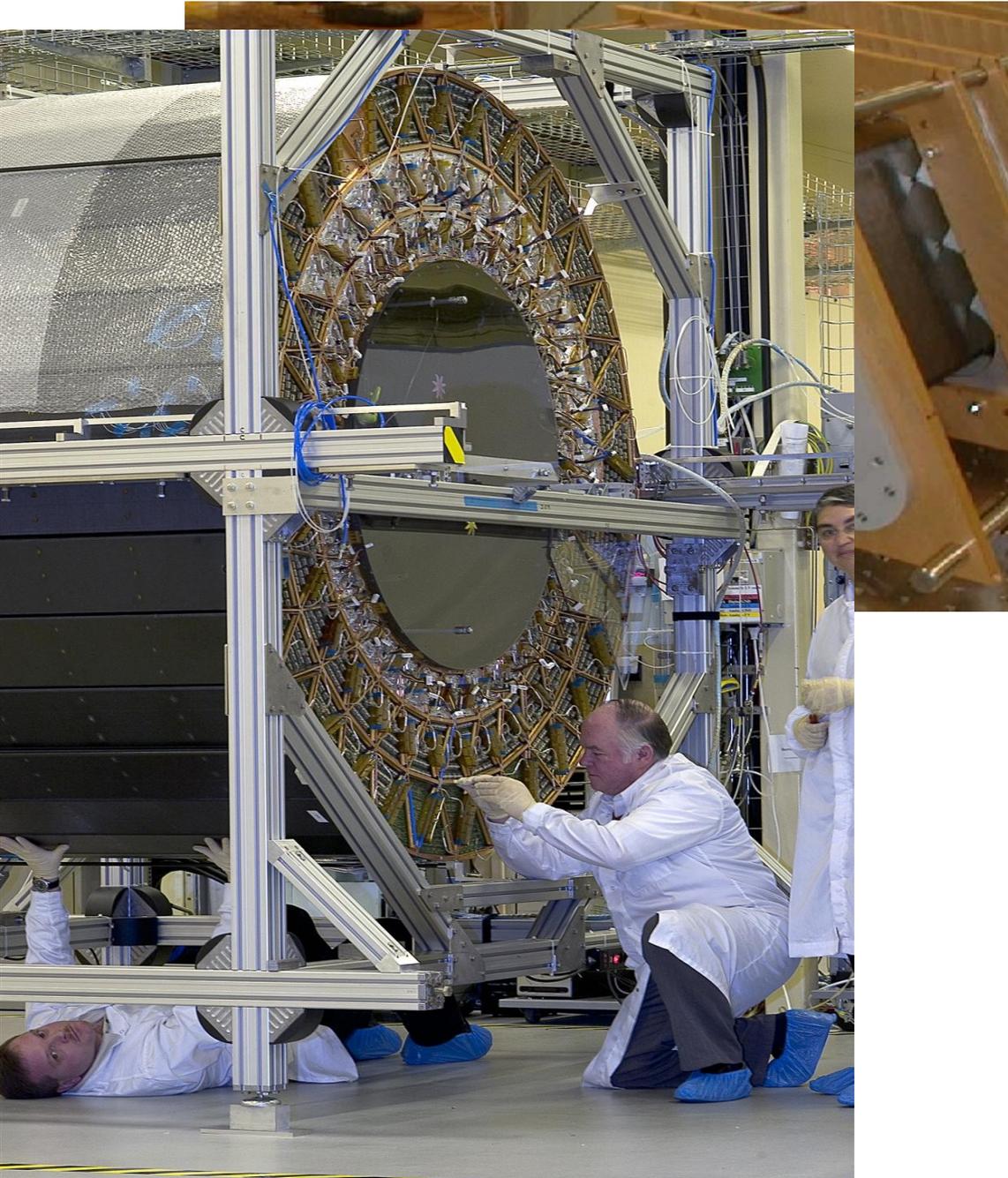


Construction highly modular



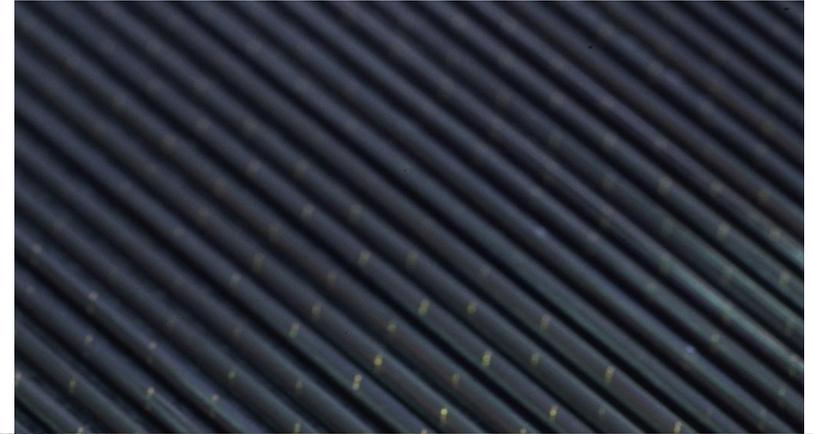
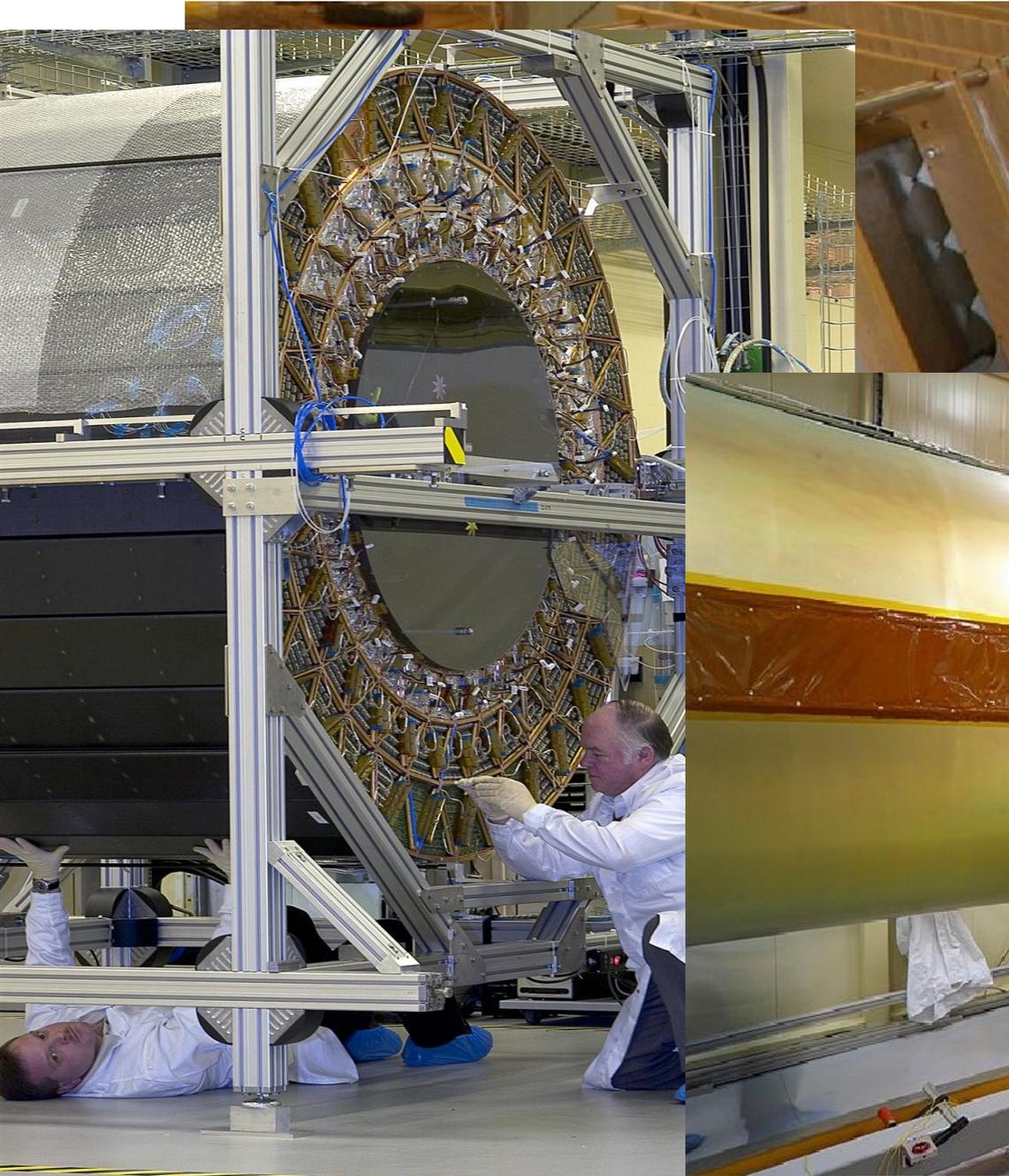


Construction highly modular

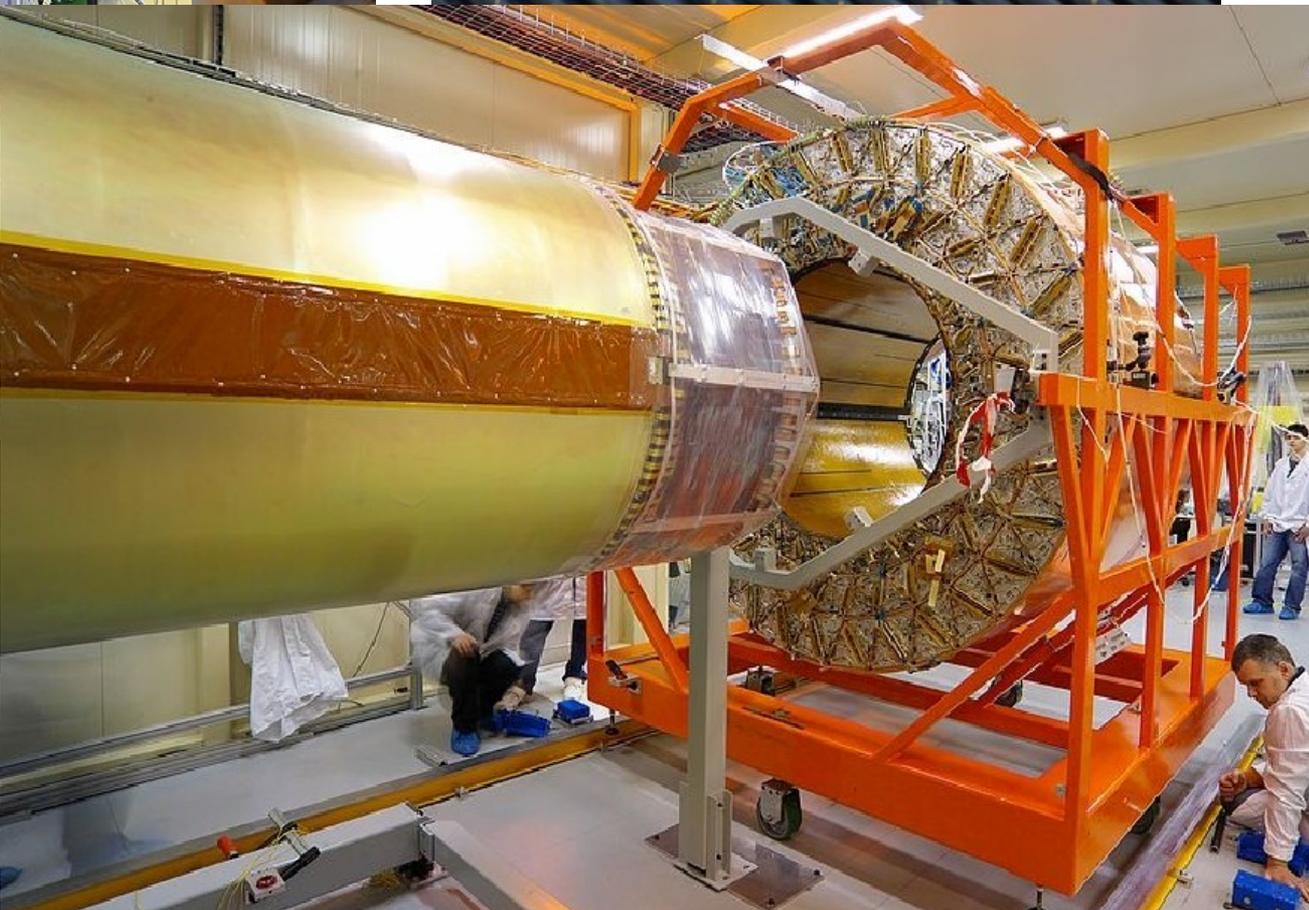
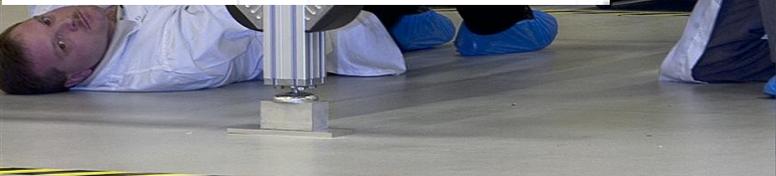
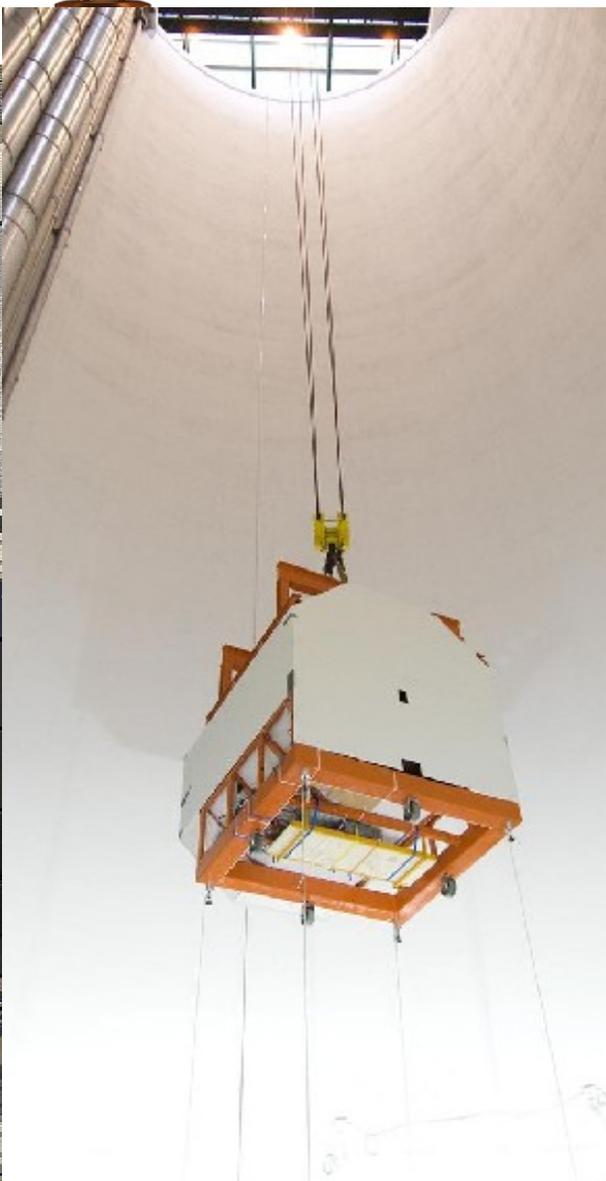




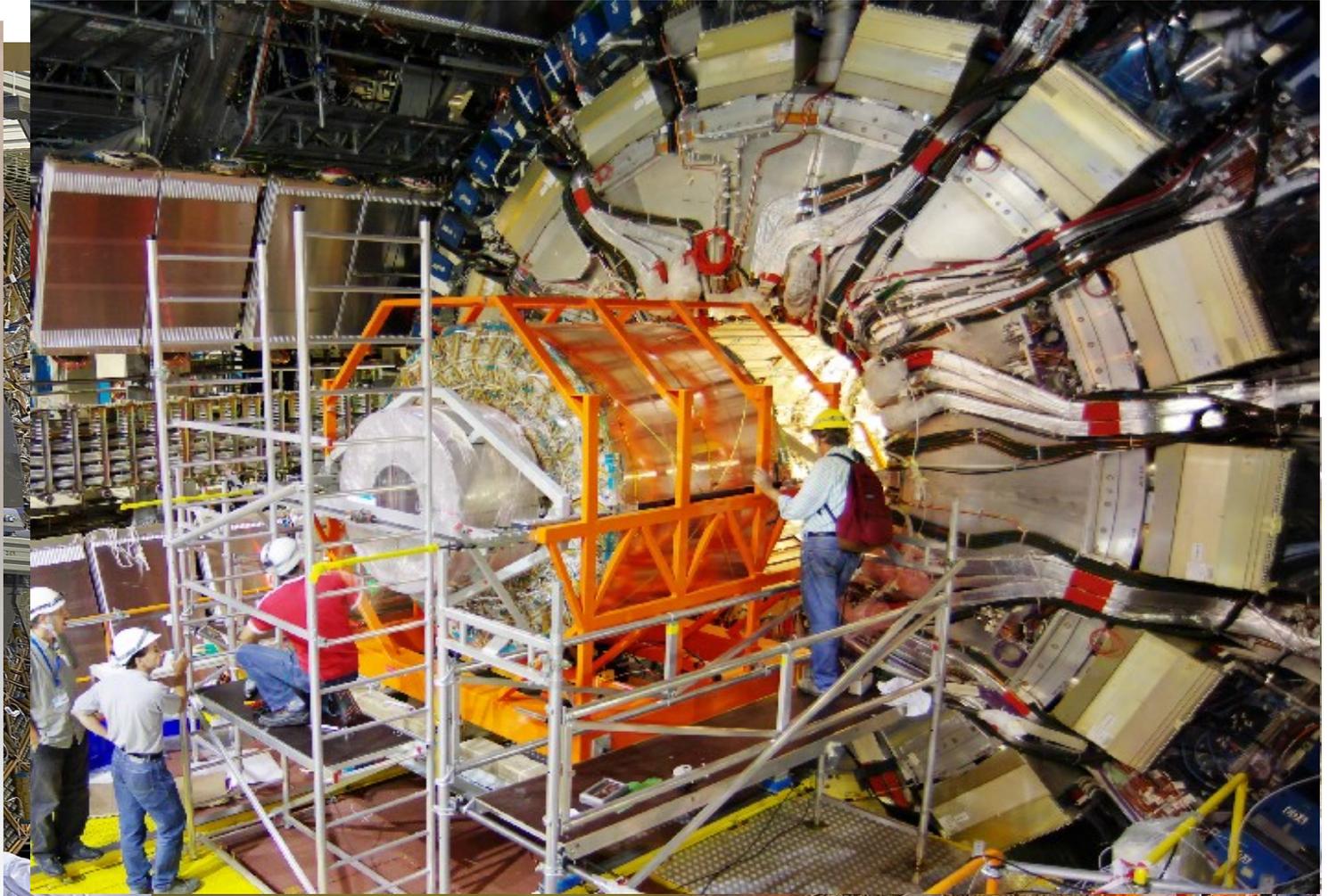
Construction highly modular



Construction highly modular



Construction highly modular

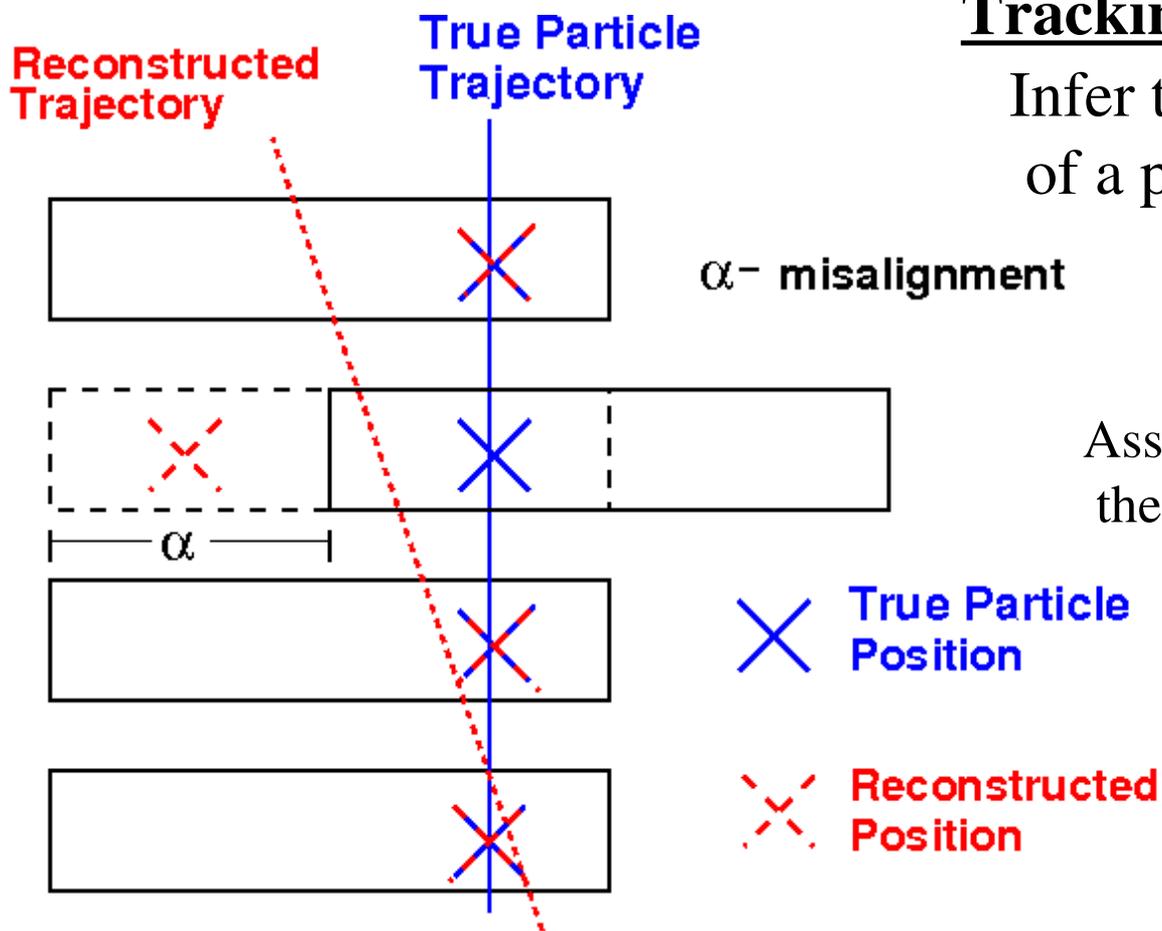




What is alignment ?



Detector positions used in reconstruction algorithms do not correspond to the actual relative positions of the installed detector.



Tracking Goal

Infer the position, direction, momentum of a particle given hits associated to it

Local Measurements

+

Assumptions about the relative location of the elements making these measurements

Why is Alignment needed?

The initial assumptions wrong



Alignment in ATLAS



Each module is positioned w/ 6 degrees of freedom (x,y,z, 3 rotations):

Si : 1744 pixel modules, 4088 SCT modules

TRT: 176 modules

~ 35,000 parameters!

Different scales of mis-alignments:

Relative Sub-detector (Si / TRT , Barrel, Endcap)

- Largest impact on physics

Internal Sub-detector

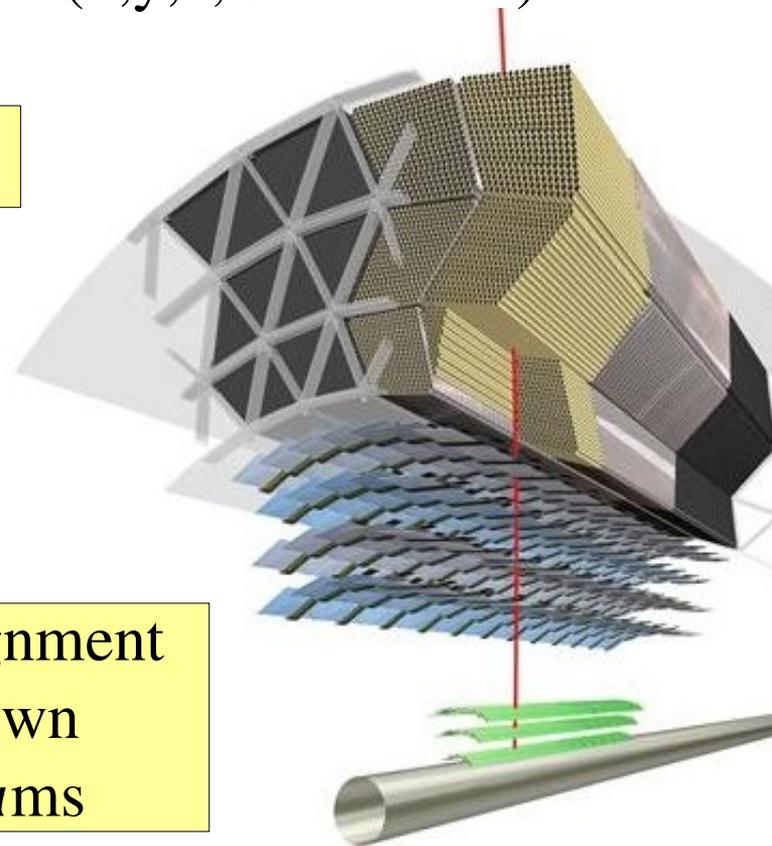
- Requires more statistics,
- Needed for ultimate precision

Requires alignment
to be known
order $10 \mu\text{ms}$

Alignment Objective

Determine relative position of in-situ detectors with the precision that alignment uncertainties contribute to less than 20% of the track parameter resolution for muons with $p_T = 100 \text{ GeV}$.

* precision physics requires the alignment to be known to O(microns) longer term goal





Track Based Alignment



Introspective

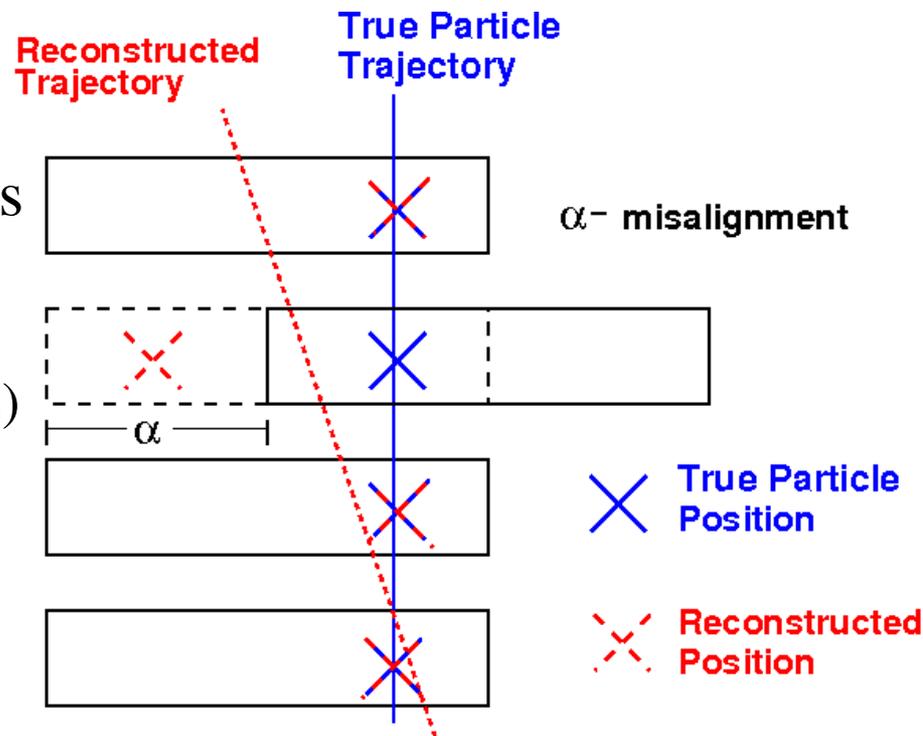
Use fact that detector misalignments affect track parameter to measure the misalignments

Define a statistic sensitive to mis-alignments
(ie: local measurements & our offline assumptions)

$$\chi^2 = \sum_{\text{hits } i} \left(\frac{m_i - r_i(\alpha)}{\sigma_i} \right)^2$$

Key properties of χ^2 function

- its an explicit function of the alignment parameters (α)
- it has a minimum at the true values of the alignment parameters



Means we can calculate $\frac{d\chi^2}{d\alpha}$

True alignment gives $\frac{d\chi^2}{d\alpha} = 0$



“Personally, I liked the university. They gave us money and facilities, we didn't have to produce anything. You've never been out of college. You don't know what it's like out there. I've worked in the private sector. They expect results.”

- *Dr. Ray Stanz*

And now, some results.



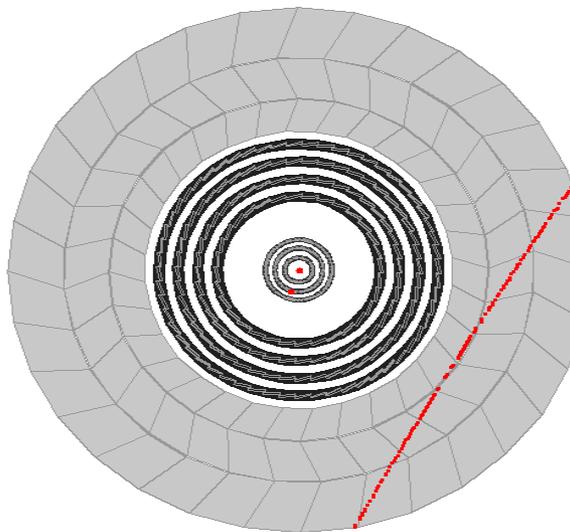
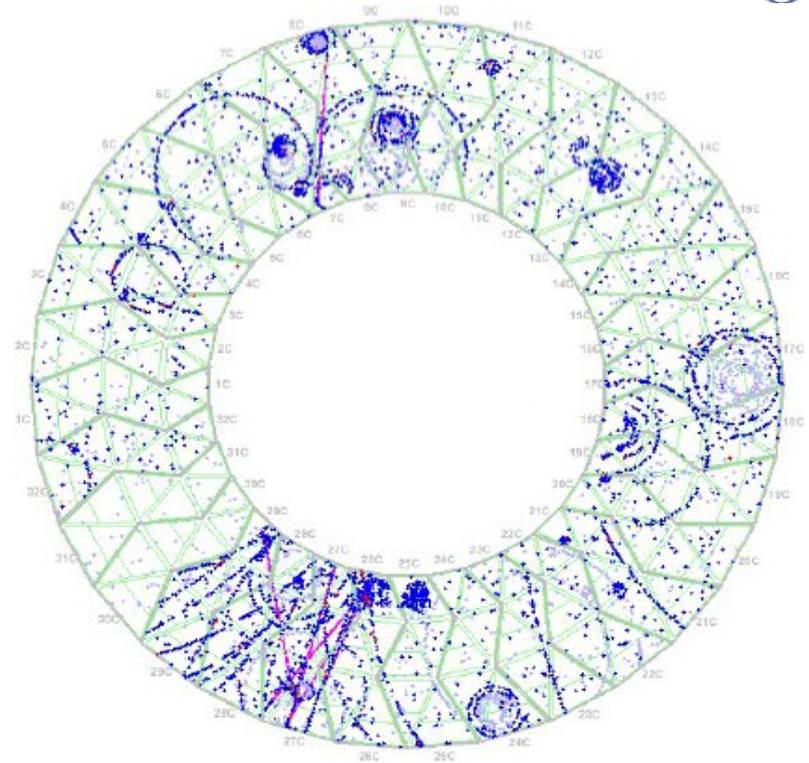
Cosmic Ray Data



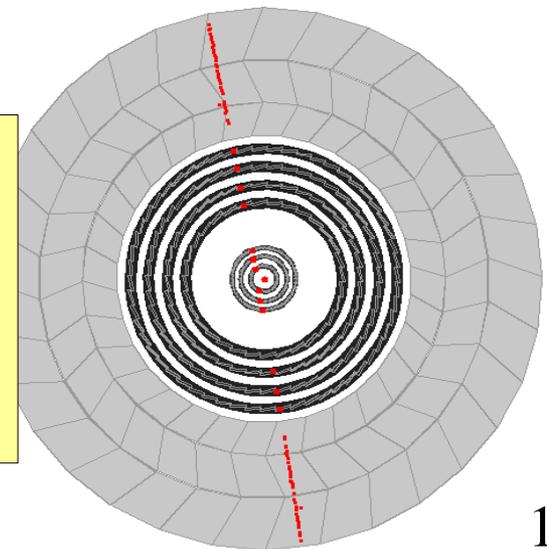
- Data taking period fall 2008.
- Over 7 million tracks reconstructed in ID
- First data available for the alignment.

Alignment has been performed:

between ID subsystems
internally with in barrel modules

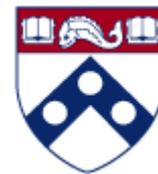


Topology of cosmic ray tracks
provides alignment algorithms
with a unique way
of seeing the detector

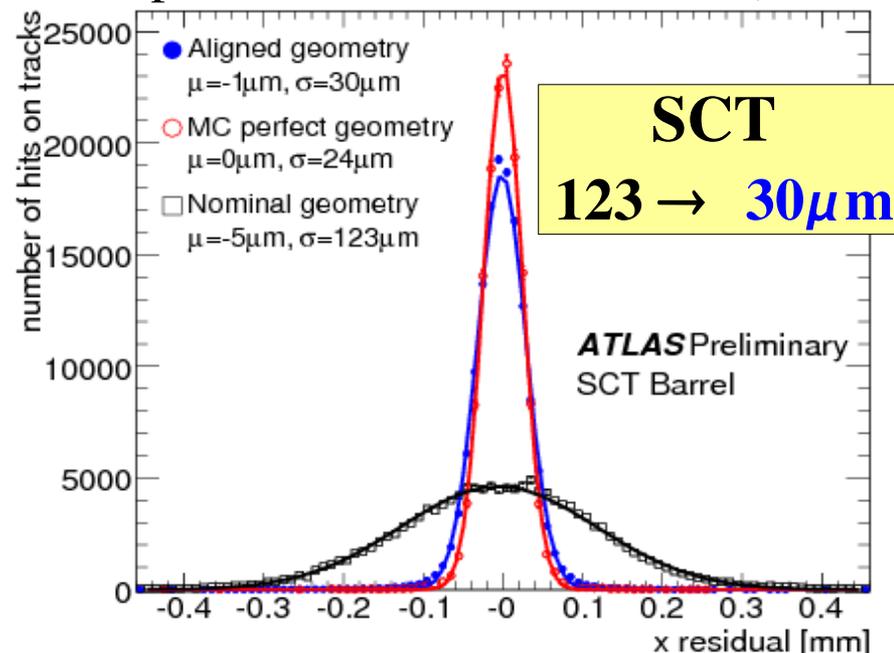
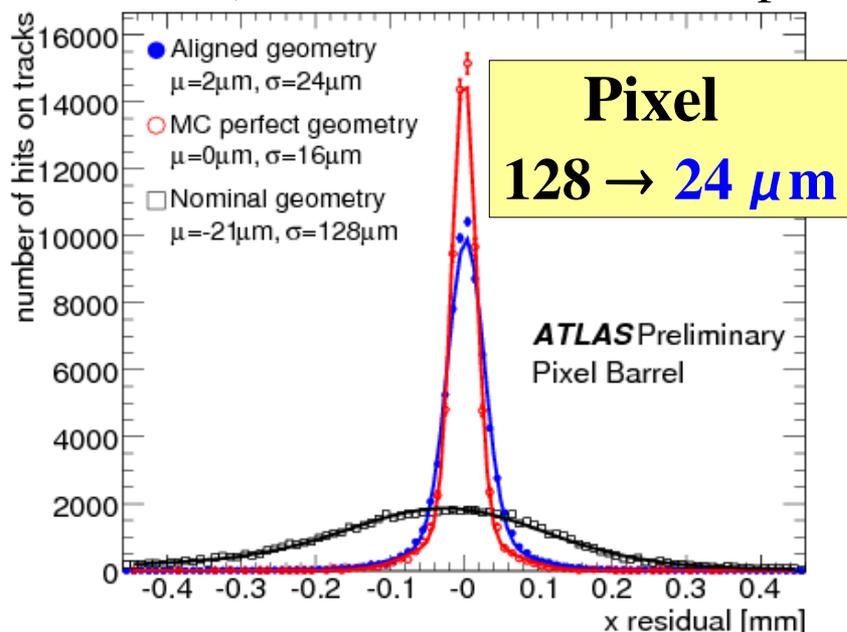




Residuals Before / After Alignment



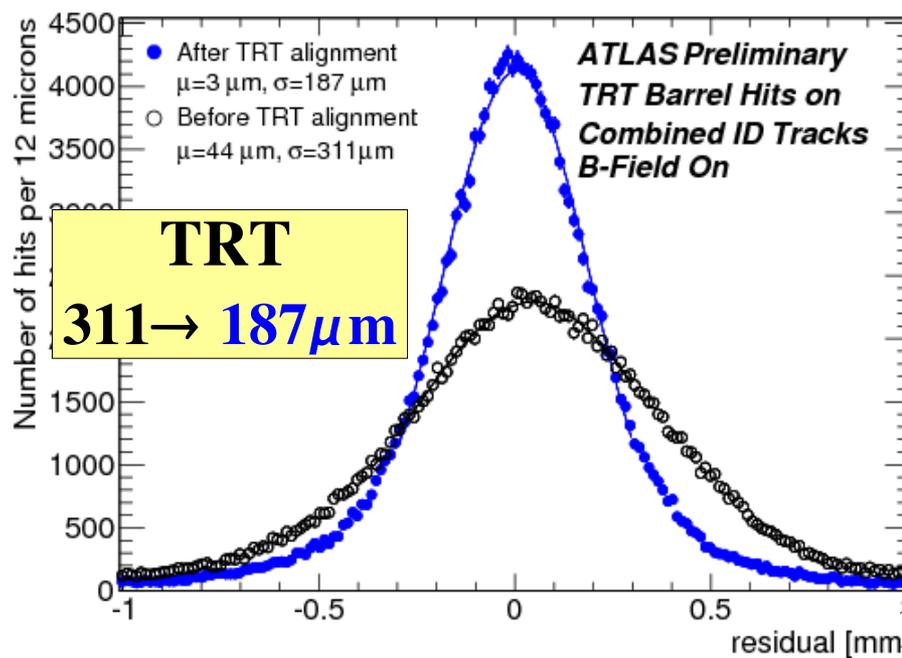
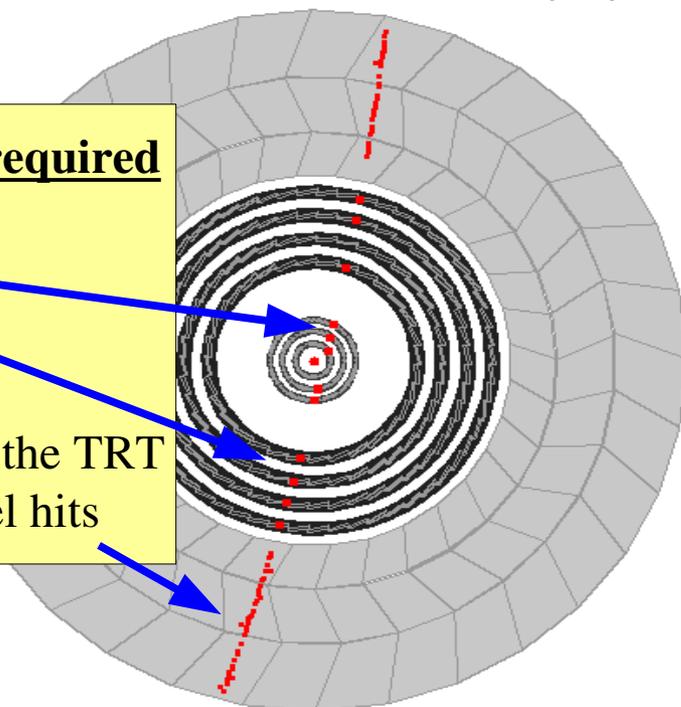
(Difference in measured position and prediction based on track fit)



Tracks shown required

- $p_T > 2 \text{ GeV}$,
- > 1 Pixel hit,
- > 8 SCT hits

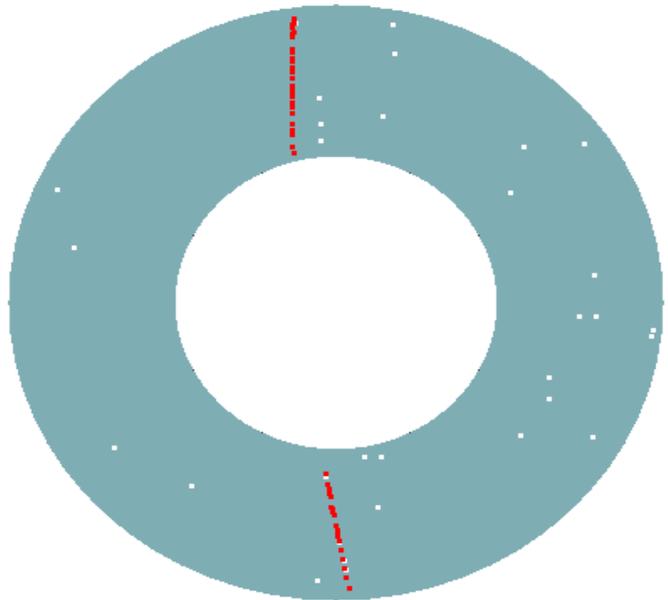
Additionally for the TRT
 > 45 TRT barrel hits



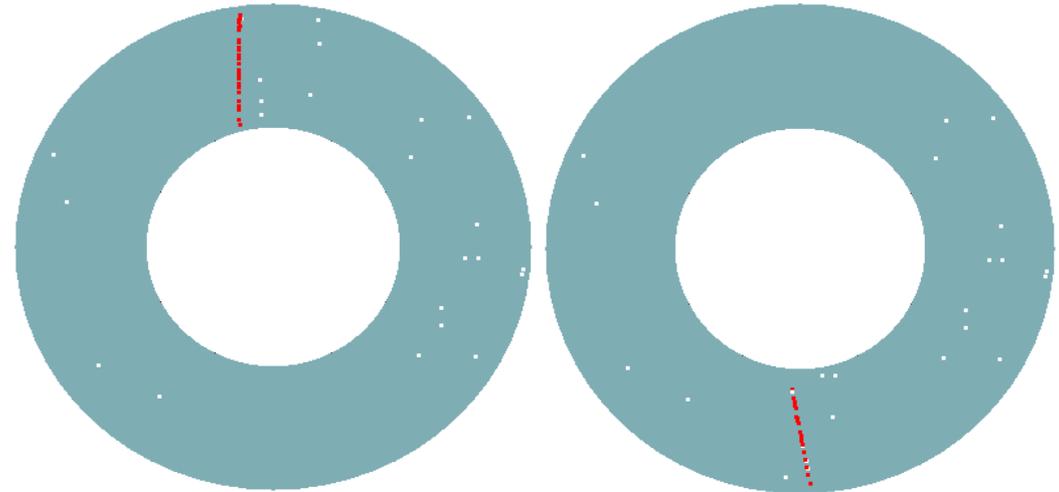


Validating alignment with data

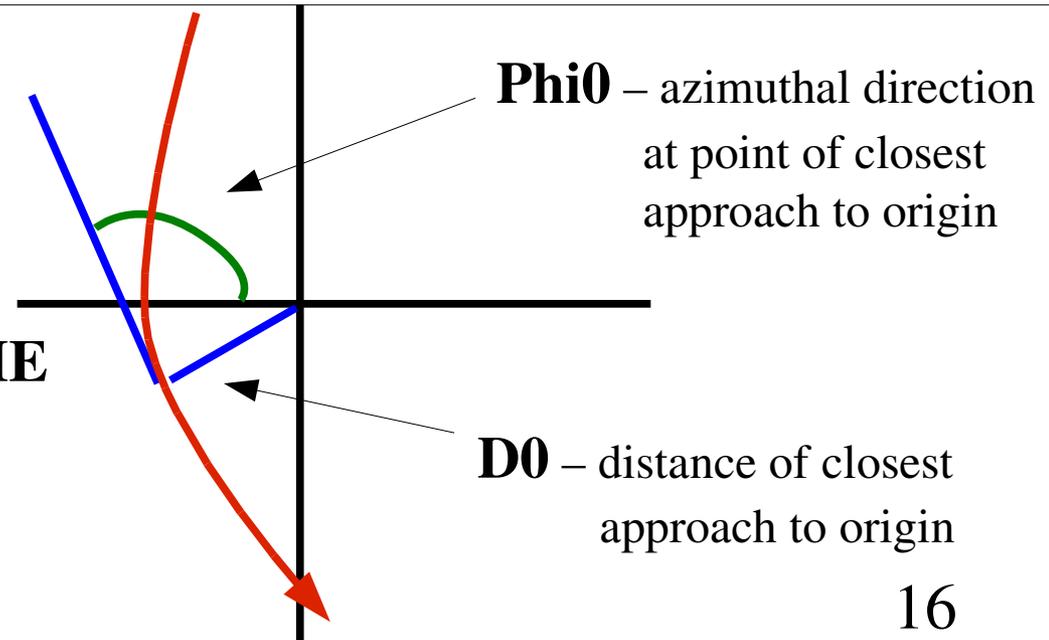
Cosmic rays come from above
traverse entire ID



Can be split in half
treated as separate particles

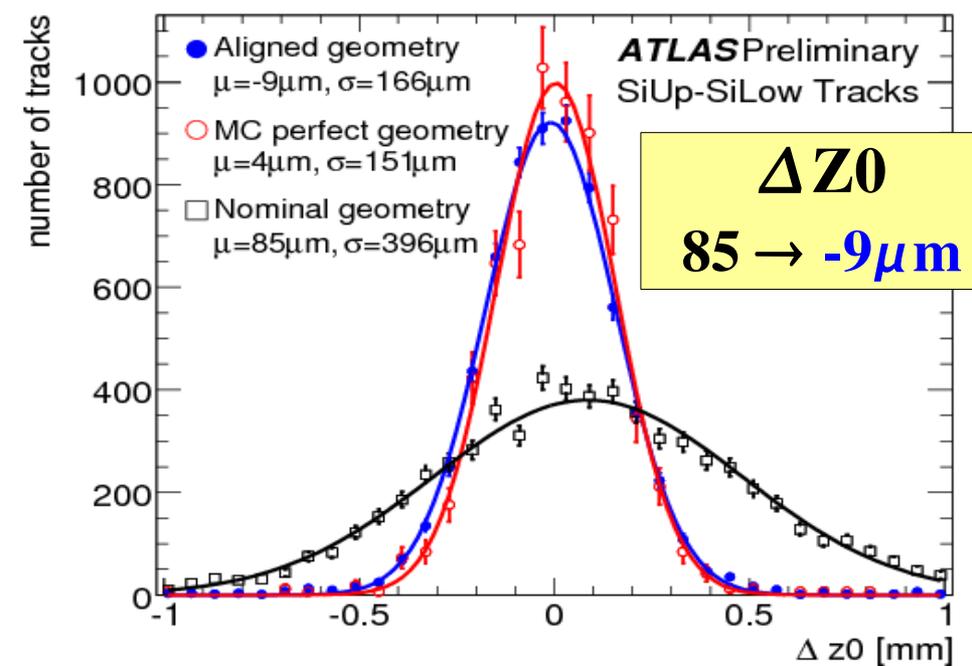
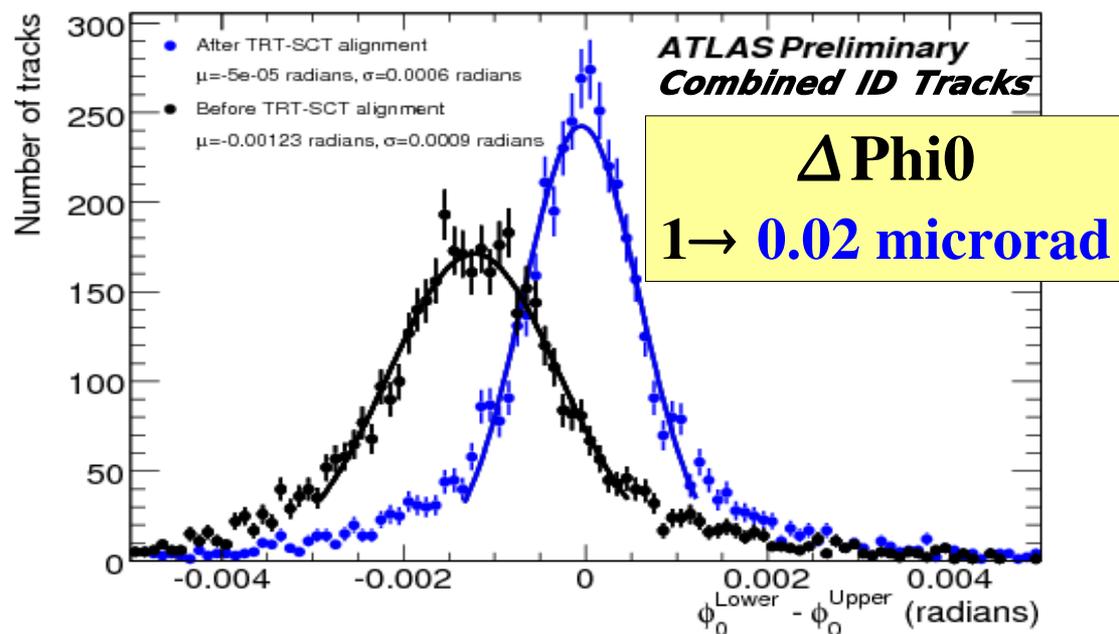
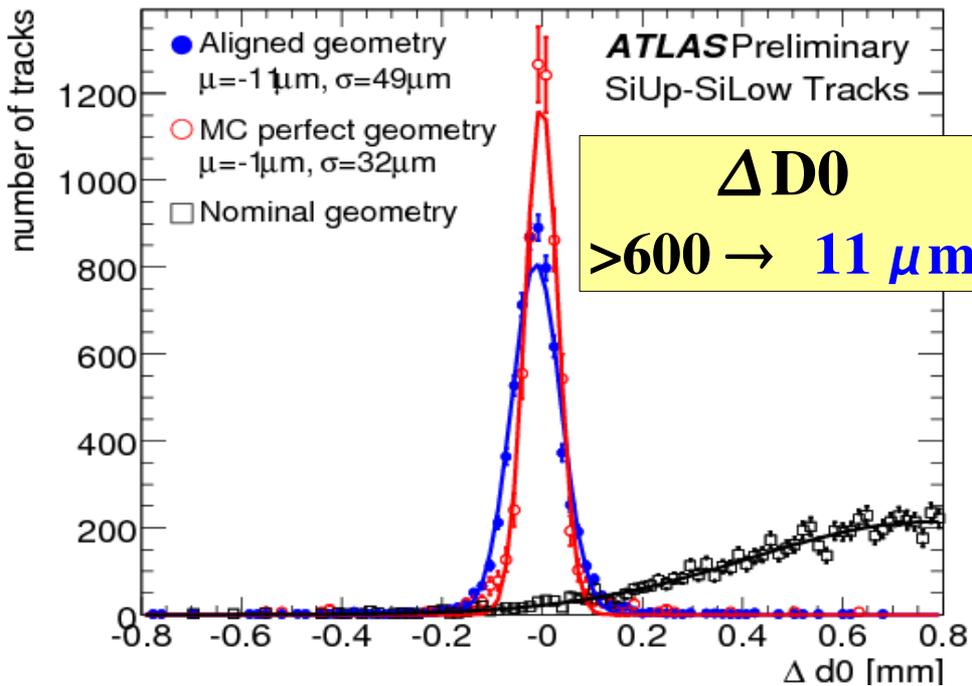
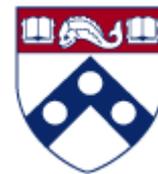


- Same particle.
- Track parameters describing the two reconstructed tracks should be the **SAME**
- Detector mis-alignment will lead to discrepancies





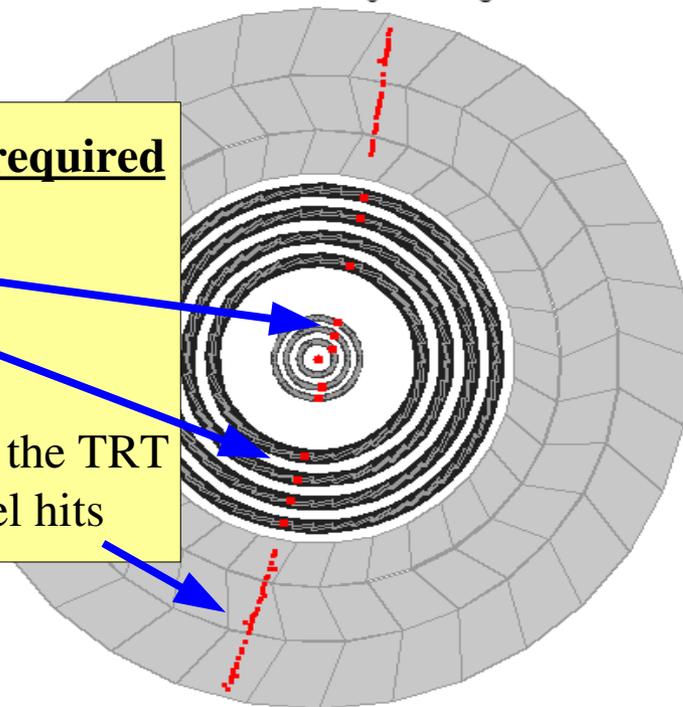
Track Parameters **Before /After** Alignment



Tracks shown required

- $p_T > 2 \text{ GeV}$,
- $> 1 \text{ Pixel hit}$,
- $> 8 \text{ SCT hits}$

Additionally for the TRT
 $> 45 \text{ TRT barrel hits}$





Conclusions

- **Overall scope and ultimate precision of ATLAS Inner Detector poses a challenging problem in terms of understanding the detector.**
Measure rotations, displacements, and distortions of over 5,000 detector elements covering volume of m^3 to 10s of microns
- **Cosmic-ray data provided huge improvements over nominal geometry and will continue to guide the way to the ultimate alignment**
- **Detector alignment procedure has been tested and validated on data from cosmic muons and is ready for collisions.**



Bonus.



Solutions to the Alignment Problem

Assembly / Survey Measurements

- External measurements of as-built detector
- after/during installation

Frequency Scanning Interferometry

- laser interference monitors differences in detector positions in real time

Track Based Alignment Algorithms

- Global χ^2
- Local χ^2
- Robust Alignment
- External constraints
 - introduction of vertex, pT, survey, e/p constraints to formalism of Global χ^2 and Local χ^2 methods

Each of these methods have been employed in solving the ATLAS ID Alignment problem to varying degrees

Will only concentrate on track based methods in the following



Track Based Alignment



Solution:

Need:

$$\frac{d\chi^2}{d\alpha} = 0$$

Approximate:

$$\frac{d\chi^2(\alpha)}{d\alpha} \approx \frac{d\chi^2(\alpha_0)}{d\alpha} + \left. \frac{d^2\chi^2}{d\alpha^2} \right|_{\alpha_0} (\alpha - \alpha_0)$$

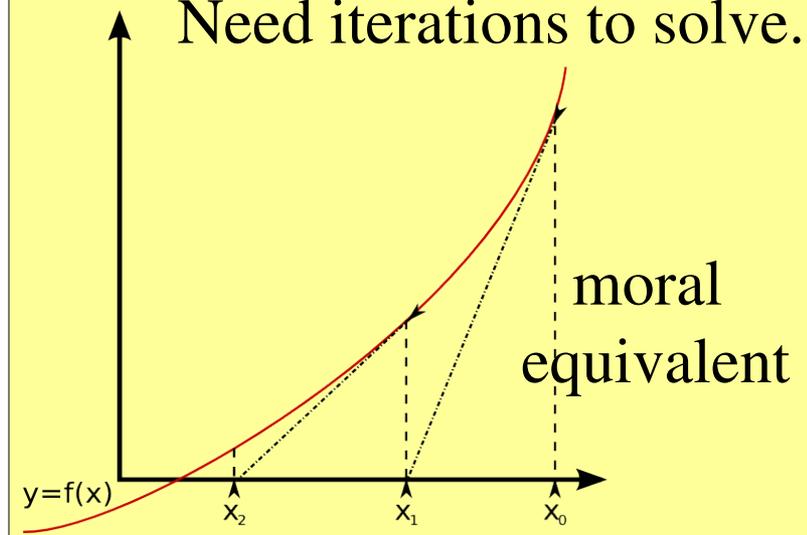
α_0 - current alignment positions

Solution:

$$\Delta\alpha \equiv \alpha - \alpha_0 = - \left(\left. \frac{d^2\chi^2}{d\alpha^2} \right|_{\alpha_0} \right)^{-1} \frac{d\chi^2(\alpha_0)}{d\alpha}$$

N x N matrix
with $N \sim 35K$
Inversion non-trivial

In general highly non-linear
Need iterations to solve.





Global Vs Local

- Described Global χ^2 method.
- Local χ^2 method exactly the same except:

$$\left(\frac{d^2 \chi^2}{d\alpha^2} \Big|_{\alpha_0} \right)$$



$$\begin{pmatrix} \frac{d\chi^2}{d\alpha_1 d\alpha_1} & \cdots & \frac{d\chi^2}{d\alpha_1 d\alpha_j} & 0 & 0 & 0 & & \\ \vdots & \ddots & \vdots & 0 & 0 & 0 & \cdots & \\ \frac{d\chi^2}{d\alpha_i d\alpha_1} & \cdots & \frac{d\chi^2}{d\alpha_i d\alpha_j} & 0 & 0 & 0 & & \\ 0 & 0 & 0 & \frac{d\chi^2}{d\beta_1 d\beta_1} & \cdots & \frac{d\chi^2}{d\beta_1 d\beta_j} & & \\ 0 & 0 & 0 & \vdots & \ddots & \vdots & & \\ 0 & 0 & 0 & \frac{d\chi^2}{d\beta_i d\beta_1} & \cdots & \frac{d\chi^2}{d\beta_i d\beta_j} & & \\ & & & \vdots & & & \cdots & \end{pmatrix}$$

Pros:

- Invert smaller matrices

Cons:

- Iterations needed to handle module correlations
- Explicit information loss
- More susceptible to weak modes

α_i β_i alignment parameters for physically distinct align-able modules



Dealing with

$$\left(\frac{d^2 \chi^2}{d\alpha^2} \Big|_{\alpha_0} \right)^{-1}$$



- **Diagonalization:**

Most CPU intensive

Provides alignment parameter errors

Removal of “weak modes”

- **Full inversion:**

Still CPU intensive

Provides alignment parameter errors

- **Fast Solver Techniques**

Exploits unique properties of derivative matrix (sparseness, symmetry)

Iterative method, minimizes distance to solution

No errors provided

Done w/ $N \sim 1000$

$$\left(\frac{d^2 \chi^2}{d\alpha^2} \Big|_{\alpha_0} \right) = U D U^T$$

U – eigenvectors D - $\begin{pmatrix} \lambda_1 & & \\ & \dots & \\ & & \lambda_n \end{pmatrix}$

$$C(\alpha) = U D^{-1} U^T$$

CLHEP, LAPACK



Dealing with

$$\left(\frac{d^2 \chi^2}{d\alpha^2} \Big|_{\alpha_0} \right)^{-1}$$



- **Diagonalization:**

Most CPU intensive

Provides alignment parameter errors

Removal of “weak modes”

- **Full inversion:**

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Iterative method, minimizes distance to solution

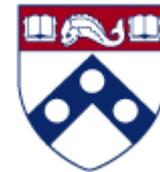
No errors provided

Done w/ $N > 10,000K$

Minimize the distance
defined as:

$$d = \left| \frac{d^2 \chi^2}{d\alpha^2} \Delta\alpha + \frac{d\chi^2}{d\alpha} \right|$$

MA27 Fortran
routine from HSL



Weak Modes

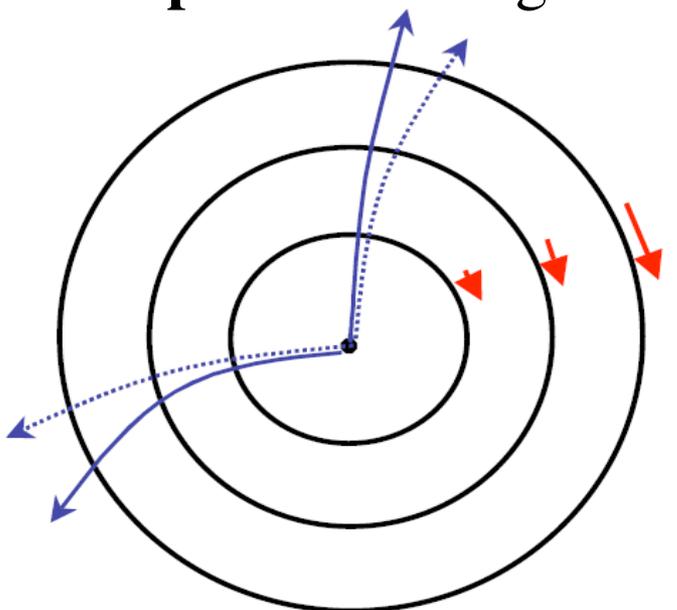
Weak Modes: Detector deformations which have little if any impact on χ^2

- inherently problematic for any method based only on χ^2
 - $\frac{d\chi^2}{d\alpha} = 0$ Solution is blind to multiple minima / classes of minima

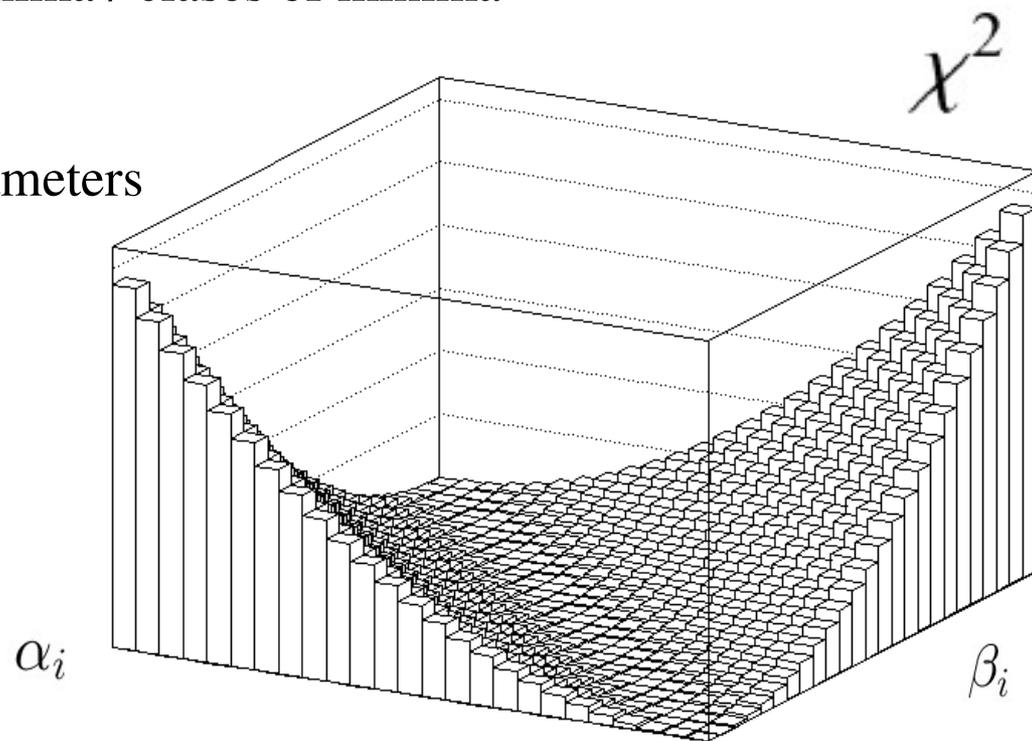
- physically important:

keep χ^2 unchanged by biasing track parameters

Example: “Clocking Effect”



pT dependent pT biasing



Whole class of systematic distortions which plague detectors with cylindrical symmetry



Dealing with Weak Modes



Weak Modes: *The real alignment problem*

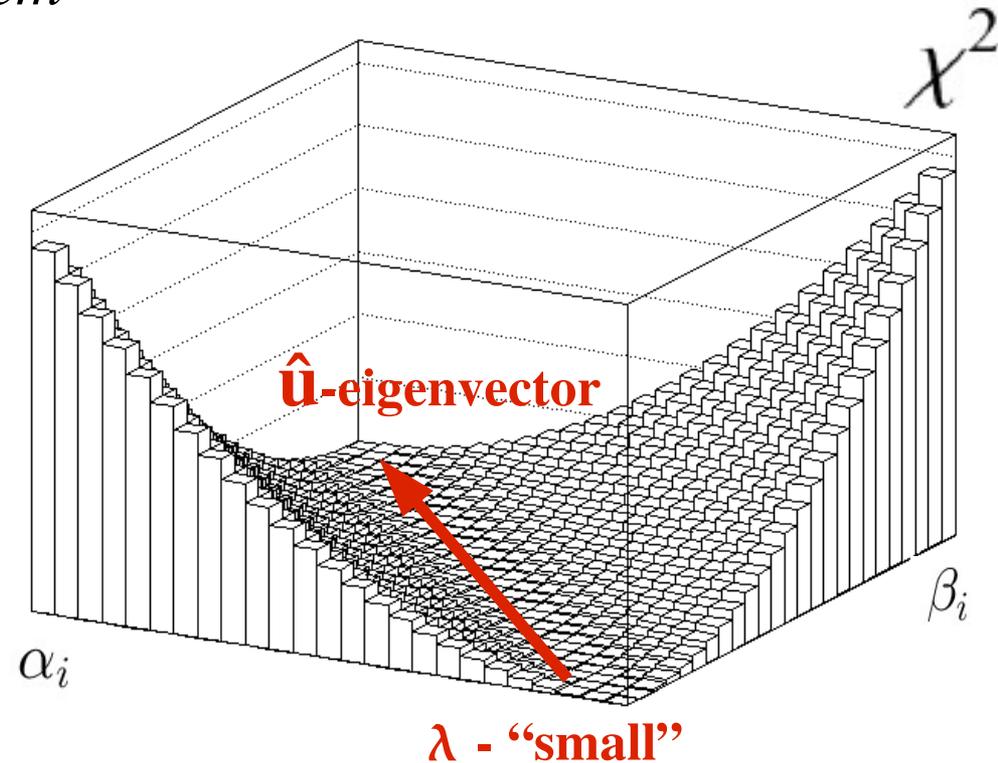
Detecting weak modes:

- Diagonalization provides means of diagnosis

$$C(\alpha) = U D^{-1} U^T$$

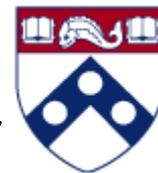
Removing weak modes:

- Explicitly remove modes below threshold
 - used to remove global movements,
 - can be dangerous / threshold arbitrary
- Enhancing definition of χ^2
 - add terms to χ^2 which depend on track parameters (eg: pT constraint, e/p)
- Event topology
 - χ^2 landscape highly dependent on event properties.
 - different events = different weak mode (a good thing!)
 - cosmic rays/beam halo, long lived decays.





Full Scale Test Alignment Procedure

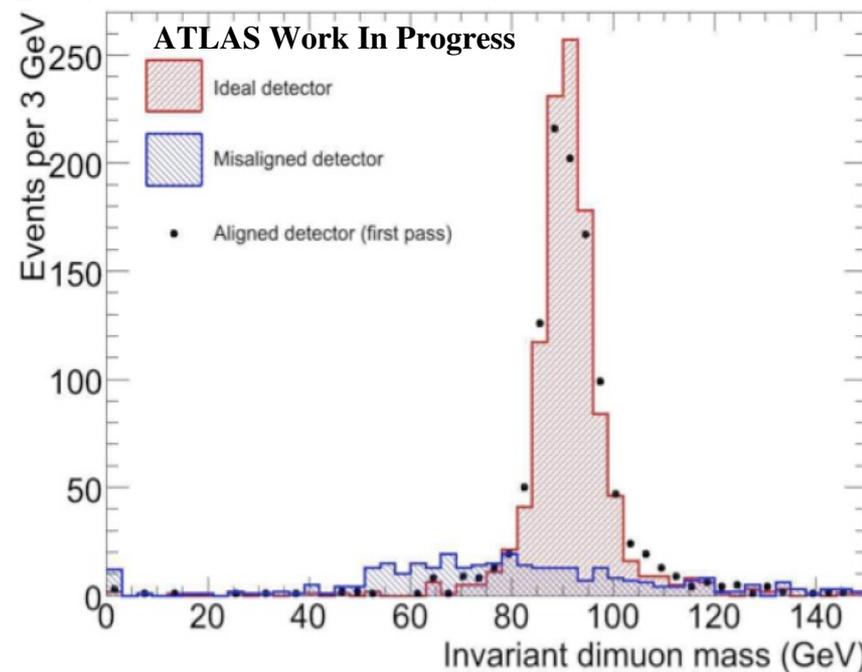


Large sample of events simulated with realistically misaligned geometry

GOAL:

- Exercise alignment algorithms, test technical infrastructure
- Provided alignment constants to the wider physics community

| Type of Mis-Alignment | Magnitude of Mis-Alignment | Number Tracks Needed |
|--------------------------------------|---|----------------------|
| Relative subsystem (Barrel / Endcap) | O(mm) translation O(mrad) rotation | 20K |
| Si Layers/Wheels | O(100 μm) translation O(0.1 mrad) rotation | 500K 50K(cosmic) |
| TRT Modules/Wheels | O(100s μm) translation O(0.1 mrad) rotation | 20K |
| Si Modules | O(< 100 μm) translation O(< 0.1 mrad) rotation | 1M |



**Muon pair mass resolution using tracks
in reconstructed in the Inner Detector**

Big Success:

Both in terms of validating the alignment procedures and in understanding problems likely to arise.



Alignment Levels



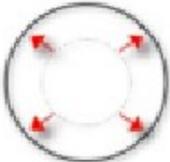
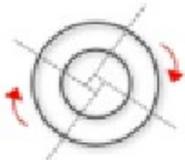
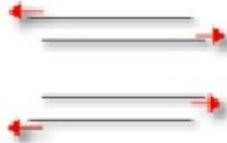
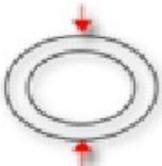
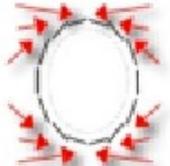
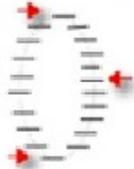
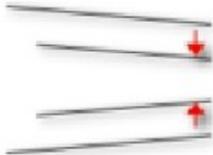
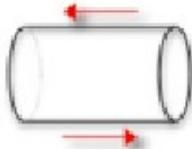
| Silicon Alignment Levels | | | | | |
|--------------------------|-------------------|------------------------------------|-------------------------|-----------------------------|-----------------------|
| Geometry Level | Structures (DoFs) | Pixel | Pixel Structures (DoFs) | SCT | SCT Structures (DoFs) |
| 1 | 4 (24) | complete pixel detector | 1 (6) | 1 barrel + 2 endcaps | 3 (18) |
| 1.5 | 7 (42) | 2 barrel half-shells + 2 endcaps | 4 (24) | 1 barrel + 2 endcaps | 3 (18) |
| 1.6 | 11 (66) | 3*2 barrel half-shells + 2 endcaps | 8 (48) | 1 barrel + 2 endcaps | 3 (18) |
| 2 | 31 (186) | 3 barrel layers + 2*3 endcap discs | 9 (54) | 4 barrel layers + 2*9 discs | 22 (132) |
| 2.1 | - (-) | - | - (-) | - | - (-) |
| 2.3 | - (-) | - | - (-) | - | - (-) |
| 2.5 | - (-) | - | - (-) | - | - (-) |
| 3 | 5832 (34992) | 1456 barrel + 2*144 endcap | 1744 (10464) | 2112 barrel + 2*988 endcap | 4088 (24528) |

| TRT Alignment Levels | | | |
|----------------------|---|---|--|
| Geometry Level | TRT | TRT DoFs | comments |
| 1 | 1 barrel + 2 endcaps | 17 | no alignment correction around the global Z-coordinate in the barrel |
| 2 | 32*3 barrel modules+ 40*2 endcap wheels | $(32 \times 3) \times 5 \text{ Dof} + (40 \times 2) \times 6 \text{ Dof} = 960$ | |



More Weak Modes



| | ΔR | $\Delta\phi$ | ΔZ |
|--------|--|---|---|
| R | Radial Expansion (distance scale)  | Curl (Charge asymmetry)  | Telescope (COM boost)  |
| ϕ | Elliptical (vertex mass)  | Clamshell (vertex displacement)  | Skew (COM energy)  |
| Z | Bowing (COM energy)  | Twist (CP violation)  | Z expansion (distance scale)  |